

Minimizing systematics with CLONES

(Constrained Local & NEsting Environment Simulations)

Jenny Sorce
and many collaborators

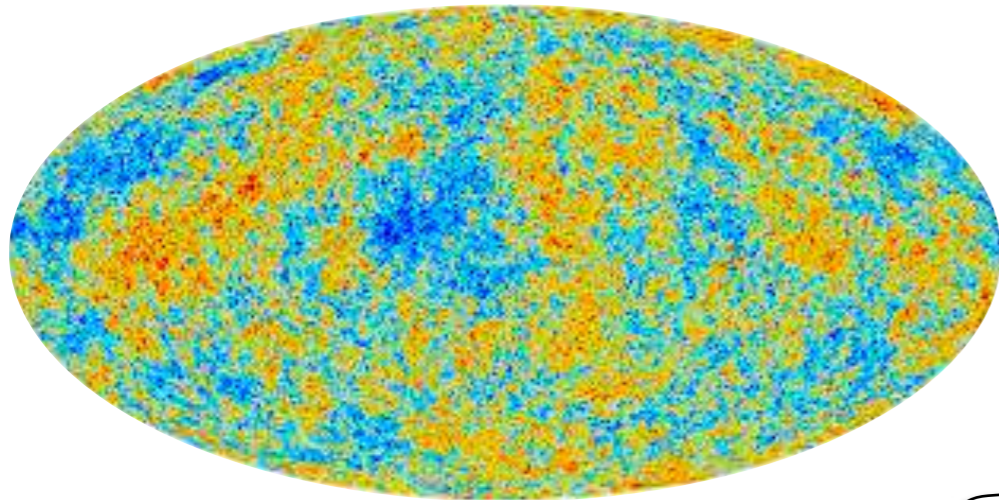
Researcher at CRIStAL, Lille & Associate Researcher at IAS, Orsay &
Guest researcher at AIP, Potsdam & CAS fellow at LMU, Munich

CosmoVerse@Lisbon - May 30th, 2023

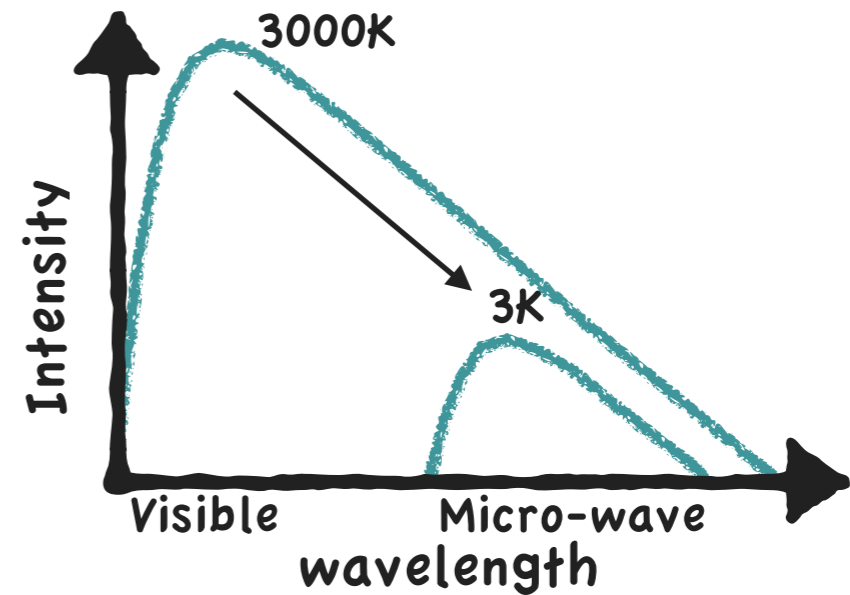


Cosmology: Λ CDM?

Cosmic Microwave Background

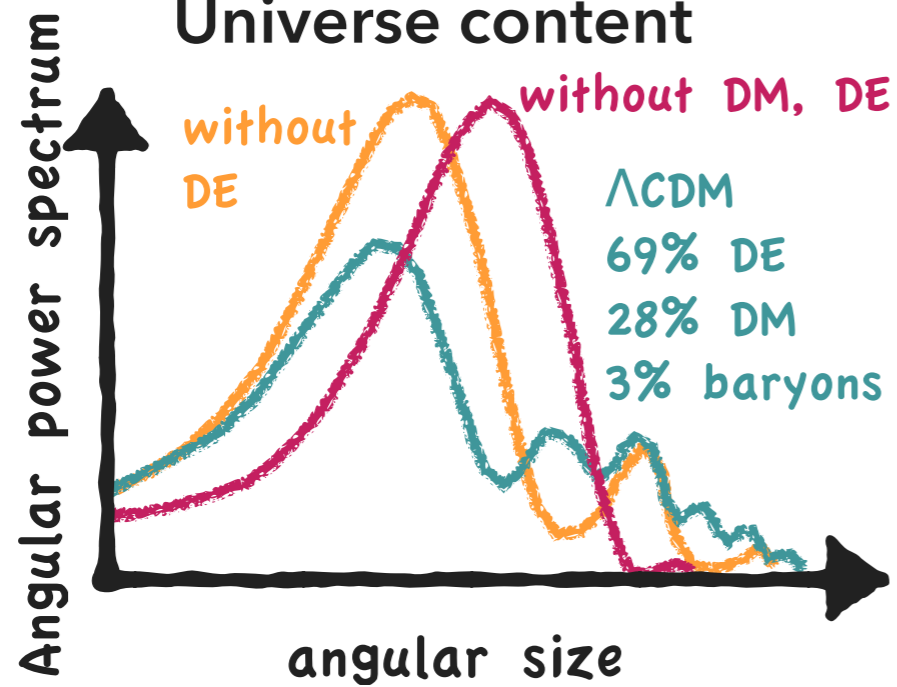


Universe expansion

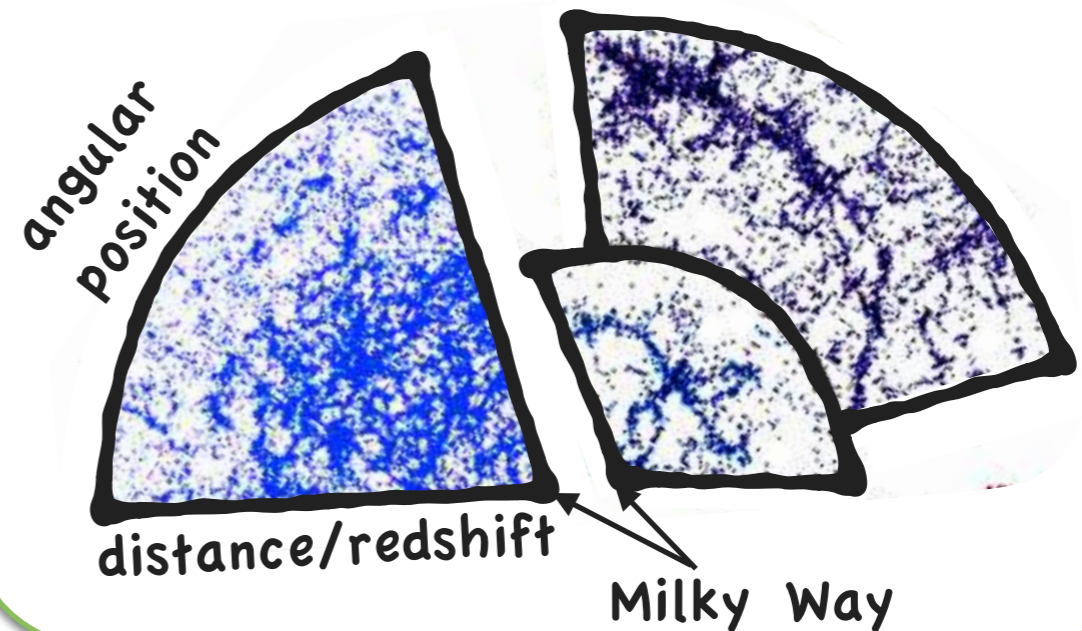


Λ CDM

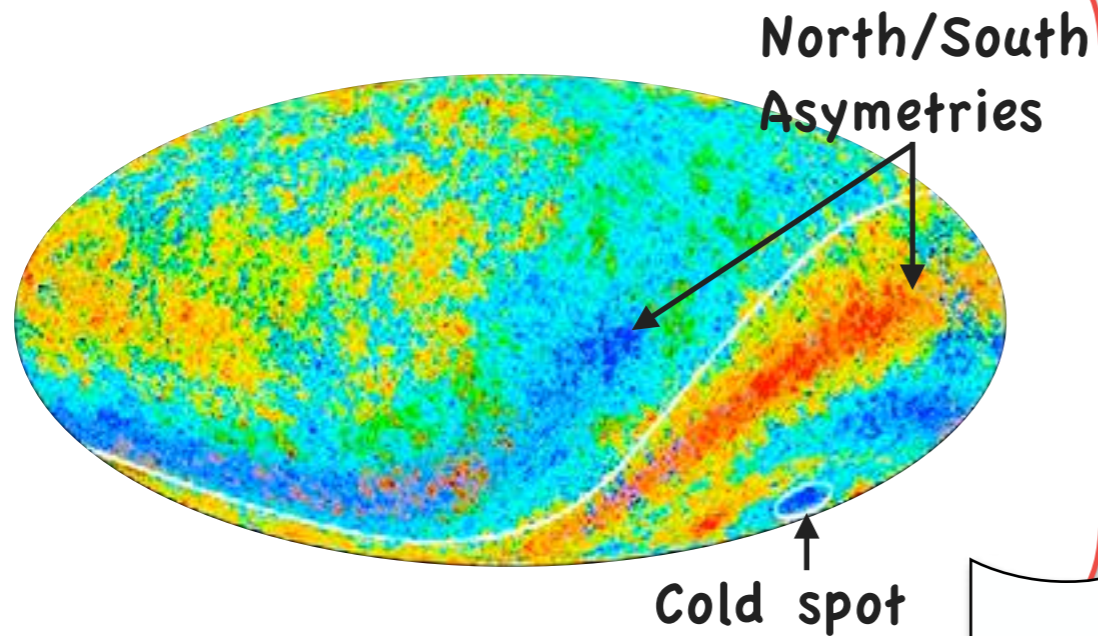
Universe content



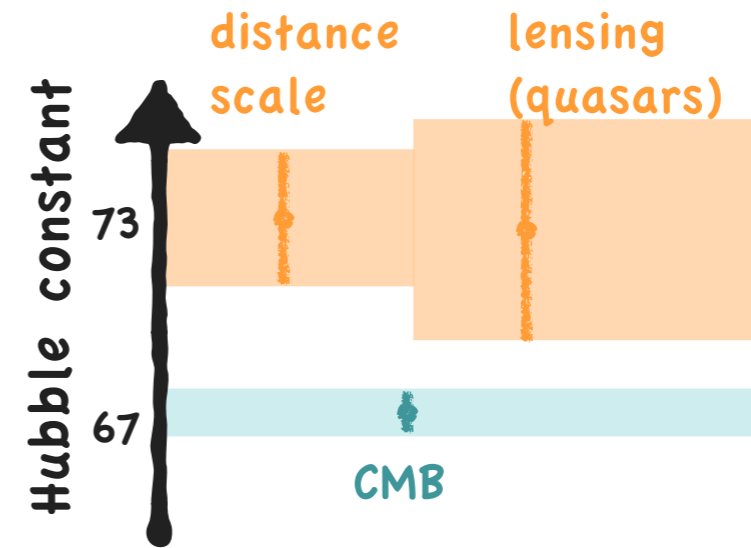
Cosmic Web and galaxies



Anomalies in the CMB

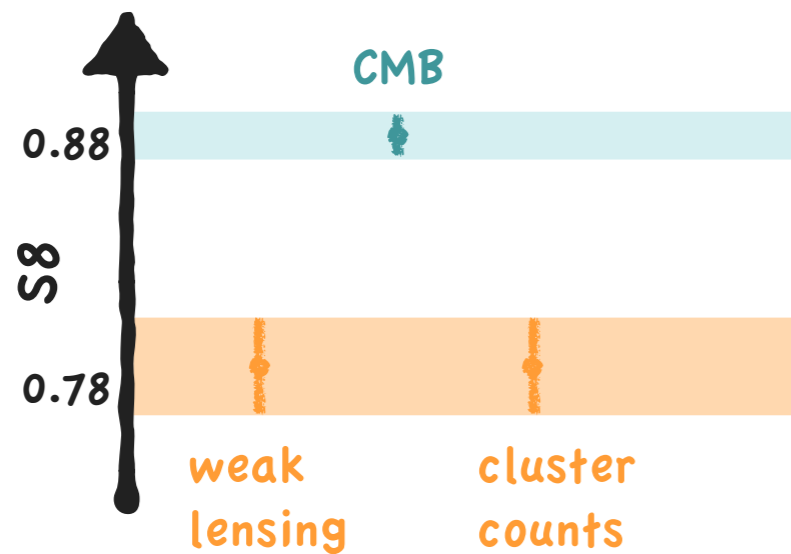


Universe expansion rate (H_0)

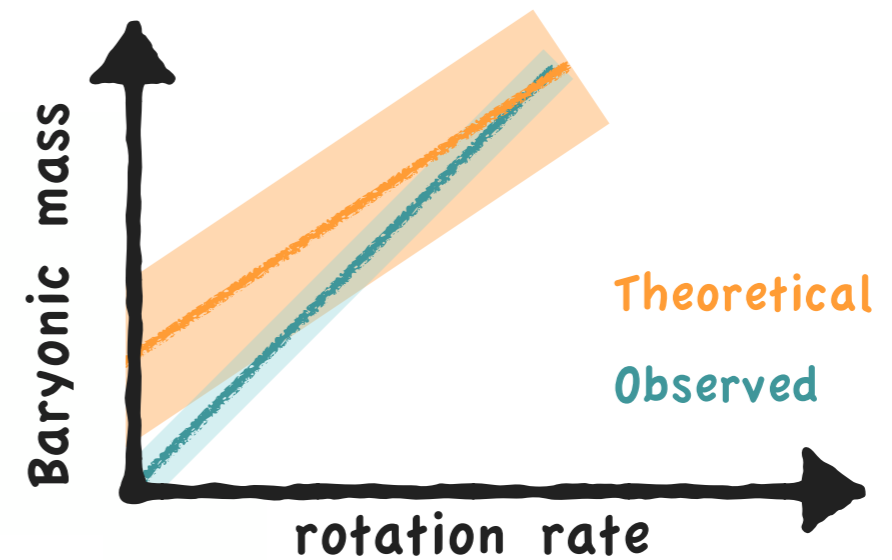


~~**Λ CDM**~~

S8 (σ_8, Ω_m)

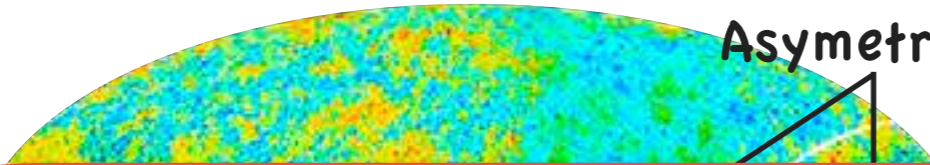


Galaxy properties



Anomalies in the CMB

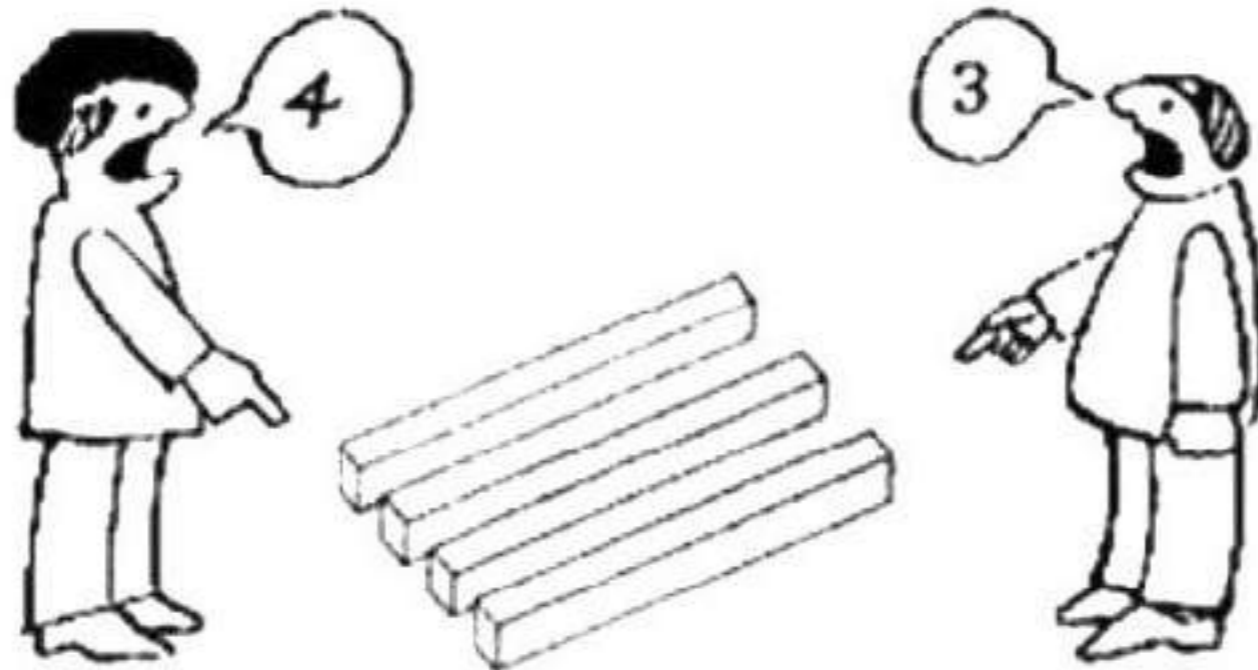
North/South
Asymmetries



Universe expansion rate (H_0)



New physics or biases/systematics?

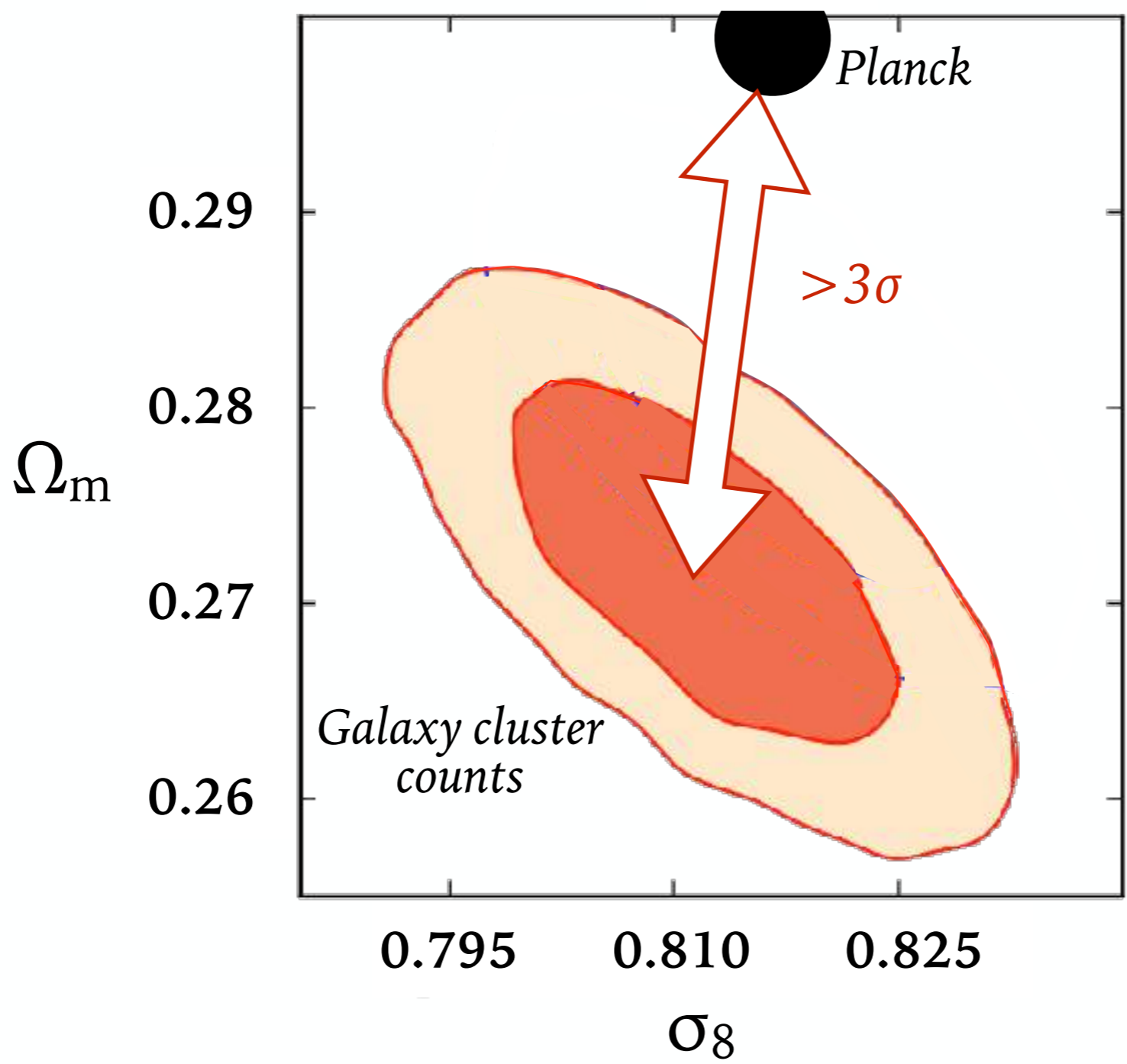


weak
lensing

cluster
counts

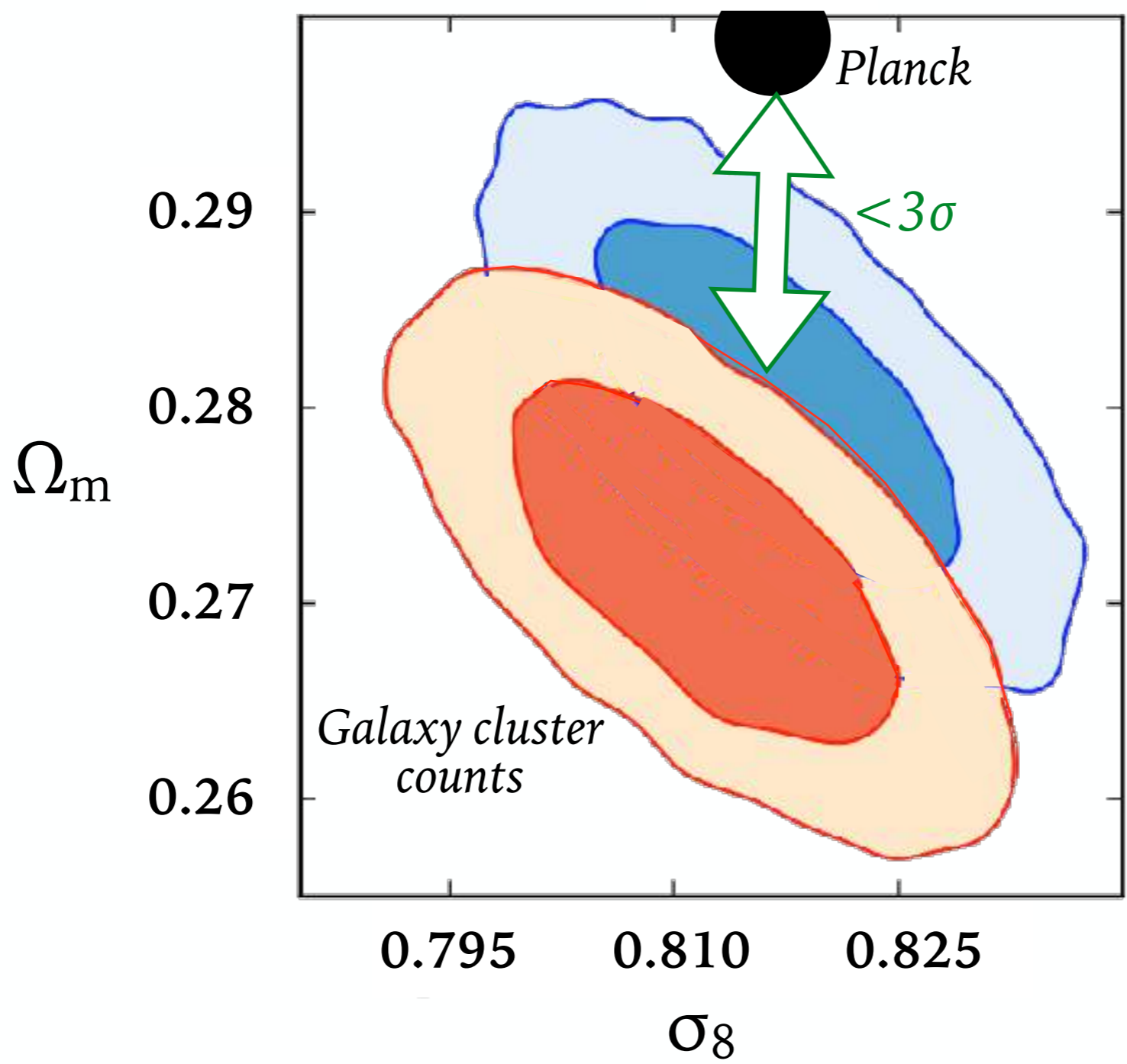


Example of S8 (σ_8, Ω_m) \blacktriangleright Galaxy cluster mass function



Planck Collaboration, Pratt+2018

Example of S8 (σ_8, Ω_m) ▶ Changing mass calibration



Planck Collaboration, Pratt+2018

Hydrostatic equilibrium :
intracluster medium

$$\frac{dP}{dr} = - \frac{G\rho M_{HE}}{r^2}$$

Spherical symmetry + no
turbulent/magnetic pressure :

$$\Rightarrow M_{HE}(r) = - \frac{rP_{th}(r)}{G\mu m_p n_e(r)} \frac{d \ln P_{th}(r)}{d \ln r}$$

Gravitational potential well : DM +
Baryons

$$M_{tot} = M_{DM} + M_{gas} + M_{stars}$$



Hydrostatic mass bias

$$\longrightarrow M_{HE} = (1 - b)M_{tot}$$

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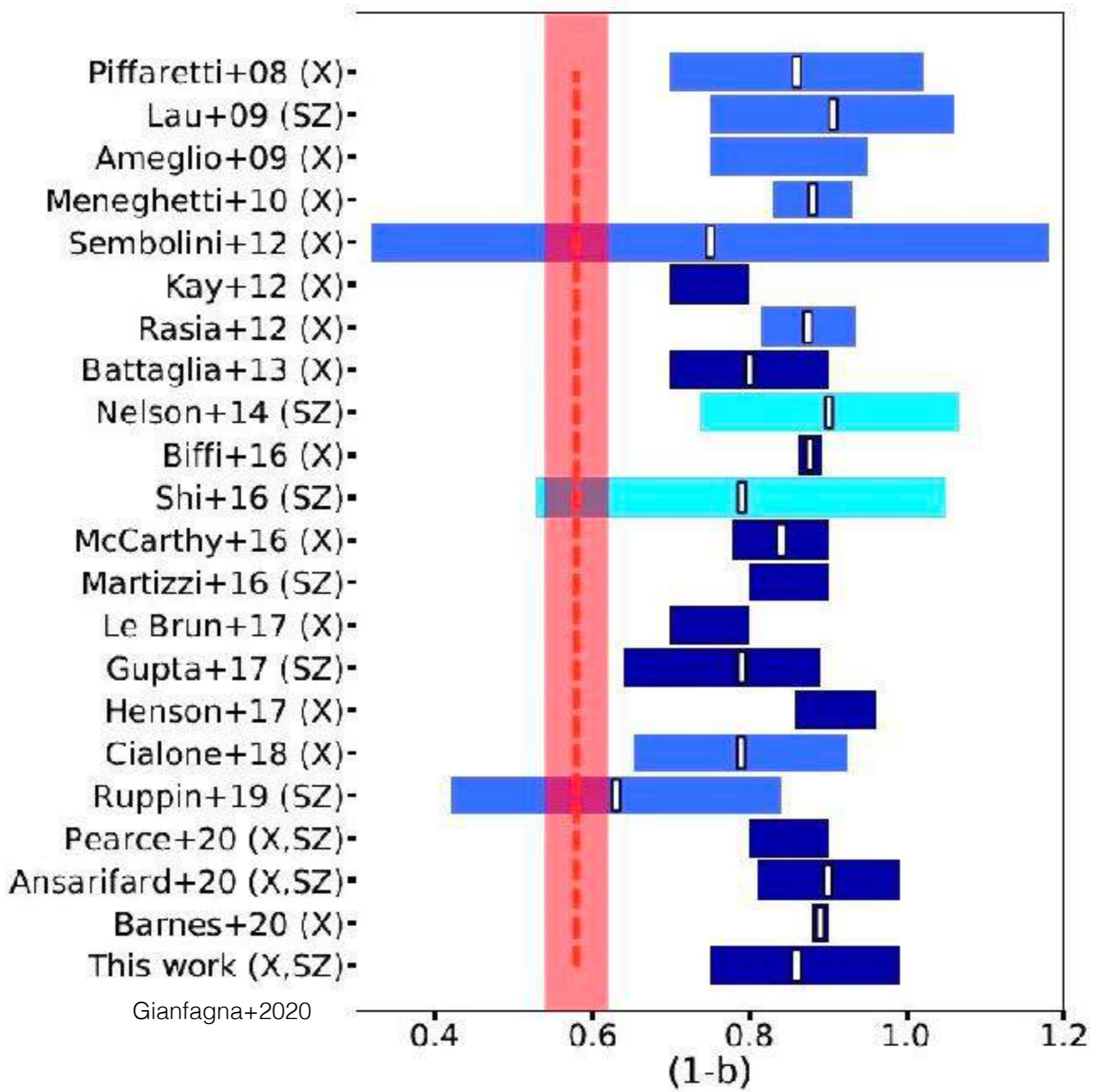


Hydrostatic mass bias

$$\longrightarrow M_{HE} = (1 - b) M_{tot}$$

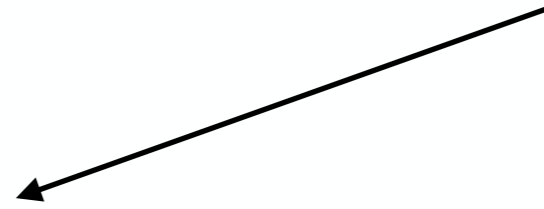
→ From cosmological simulations

Example of S8 (σ_8, Ω_m) ▶ Huge disparity



Gianfagna+2020

$$S8 (\sigma_8, \Omega_m) = X \pm \sigma_{\text{measure}} \pm \sigma_{\text{systematics}}$$

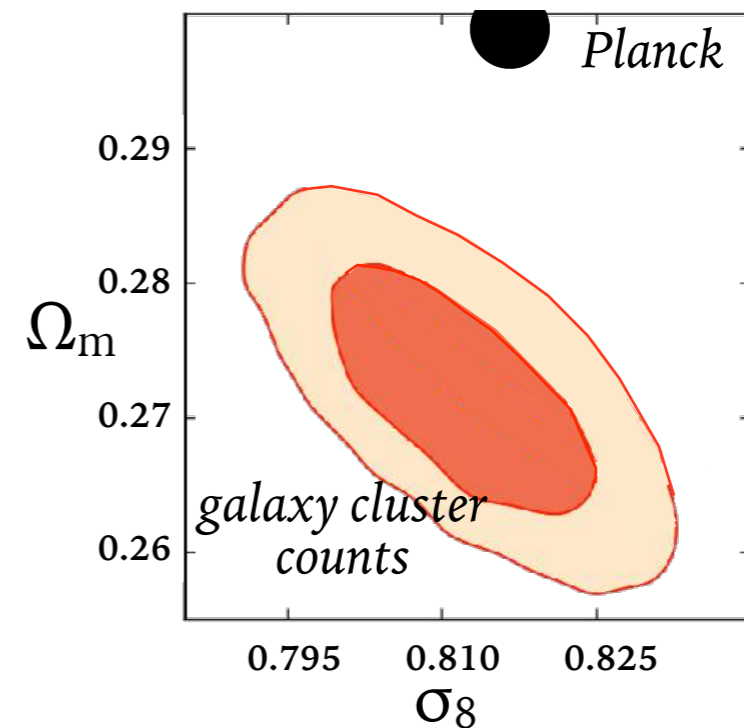
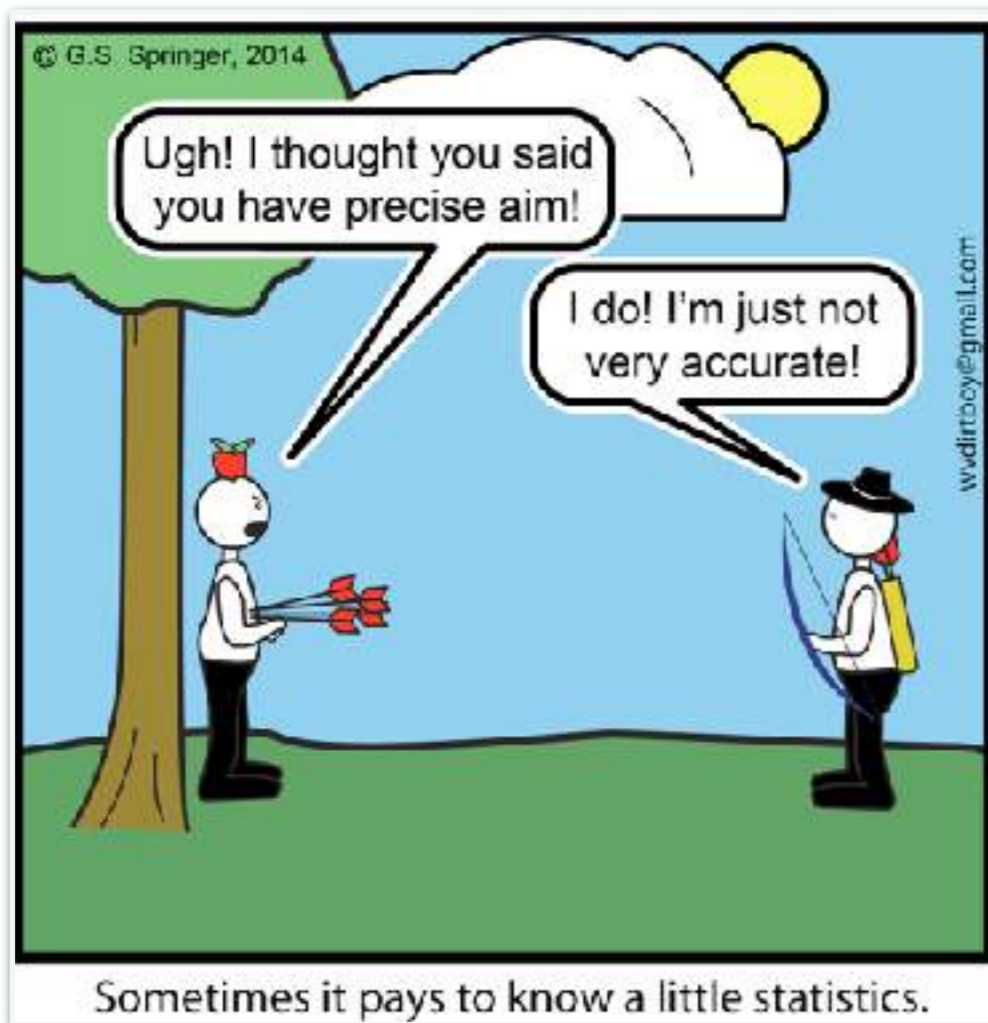


- nb measurements
- instruments/tools sensitivity
= precision

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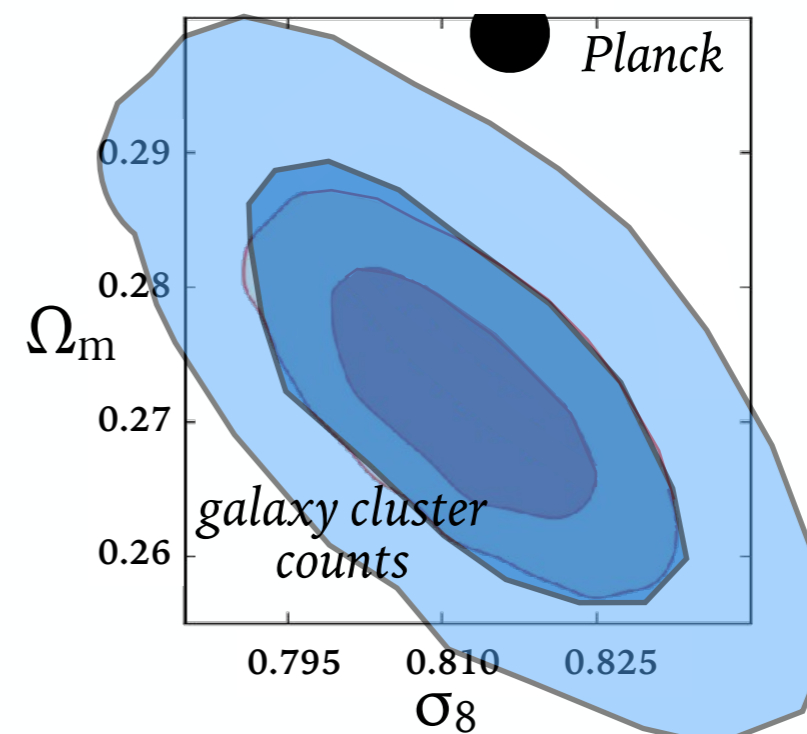
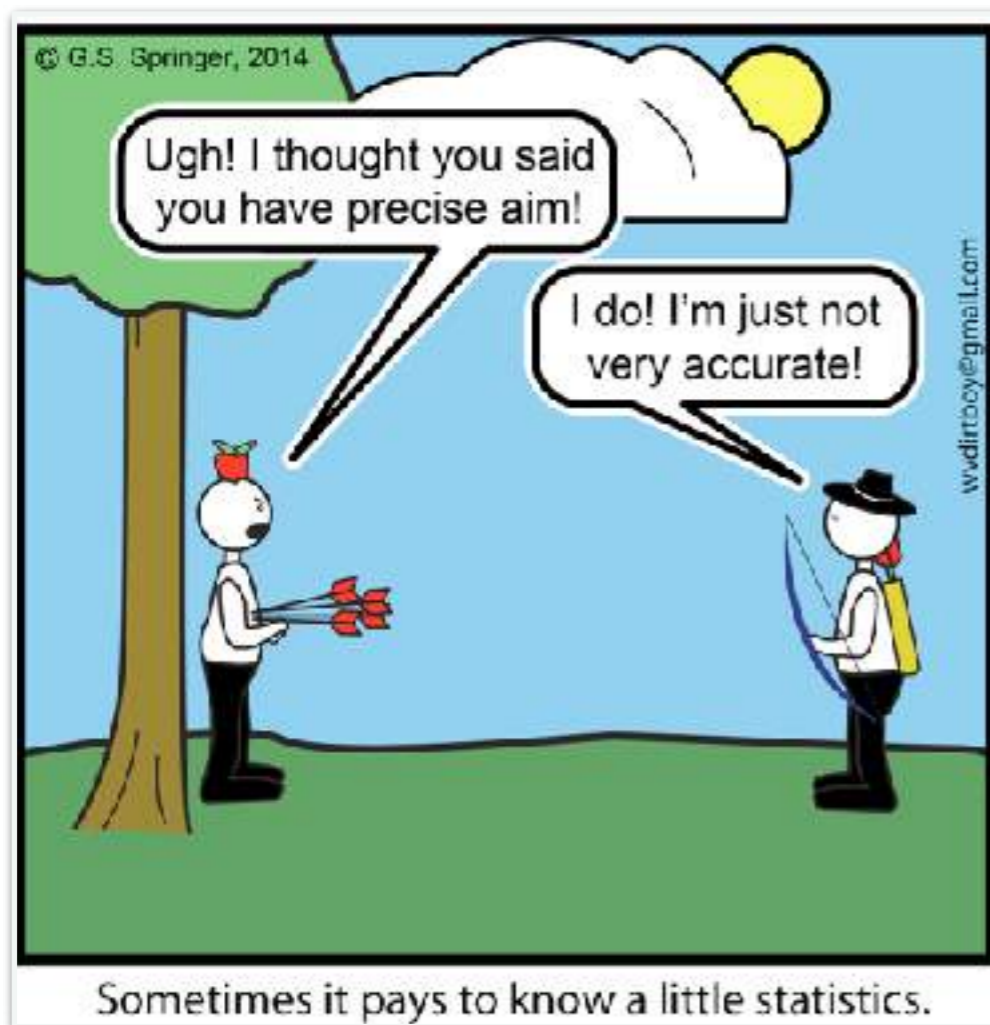
Standard cosmological simulations can give the total uncertainty but cannot reduce the systematics



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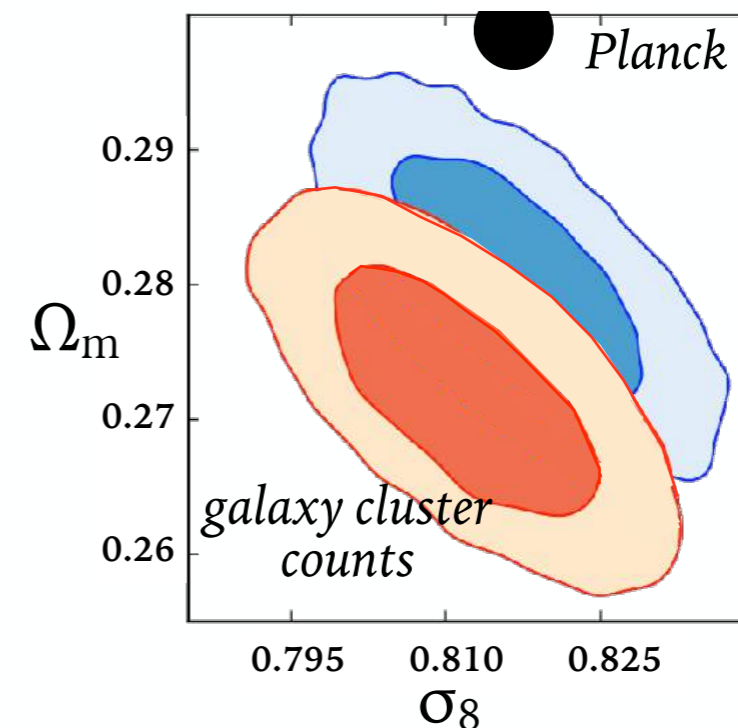
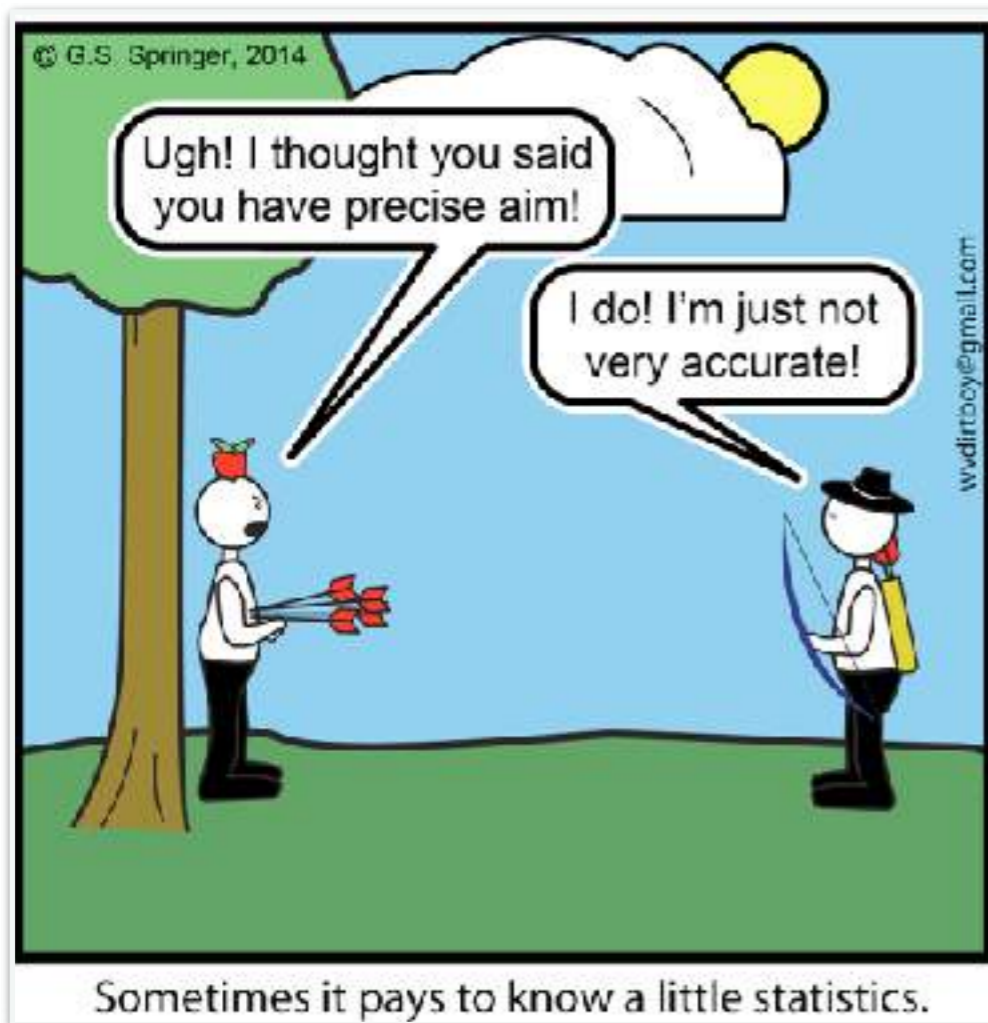
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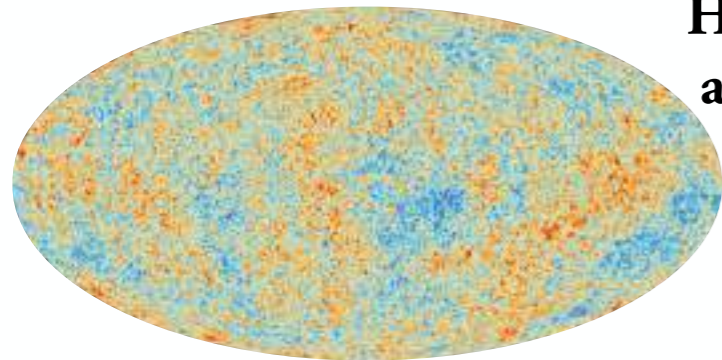


Constrained cosmological simulations can help reduce biases

Standard cosmological simulations

Initial conditions (ICs)

Part of the Universe at
13.7 light-Gyr
Photons received today
have been emitted when it
was ~380 000 yrs. old



Homogeneous
and Isotropic
Universe

→

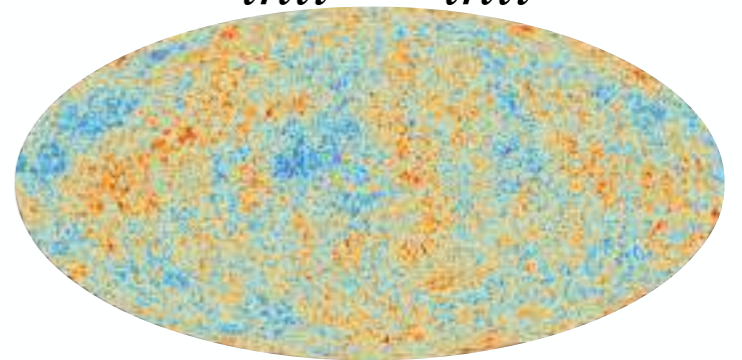
$$P(k)$$

Gaussian
initial density
field

→

$$\delta(\mathbf{k}) = \sqrt{P(\mathbf{k})} \cdot \omega(\mathbf{k})$$

→



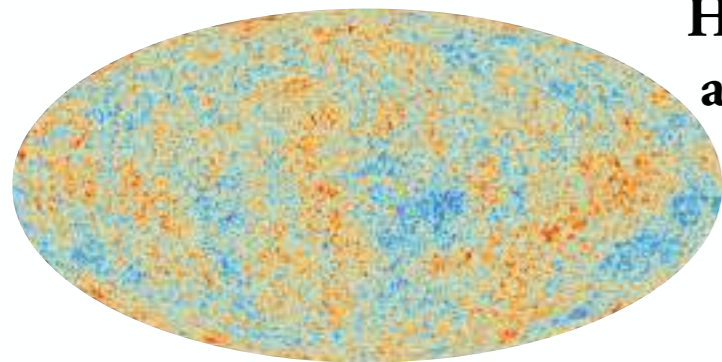
initial conditions of
a random patch of
the Universe

$$\{\delta_{init}, v_{init}\}$$

Standard cosmological simulations

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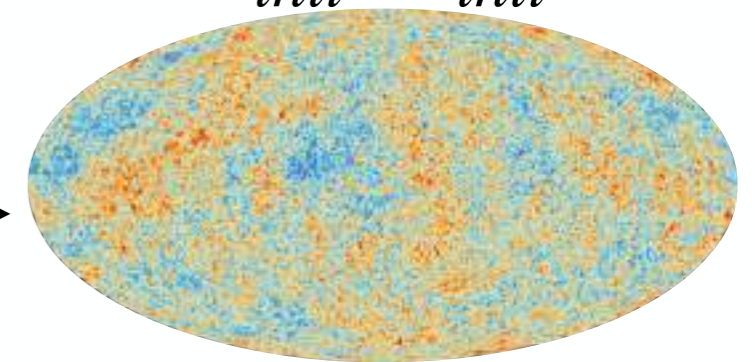
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→



initial conditions of
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the Universe

$$\{\delta_{init}, v_{init}\}$$

Linear perturbation
theory (Euler+
Continuity+Poisson)

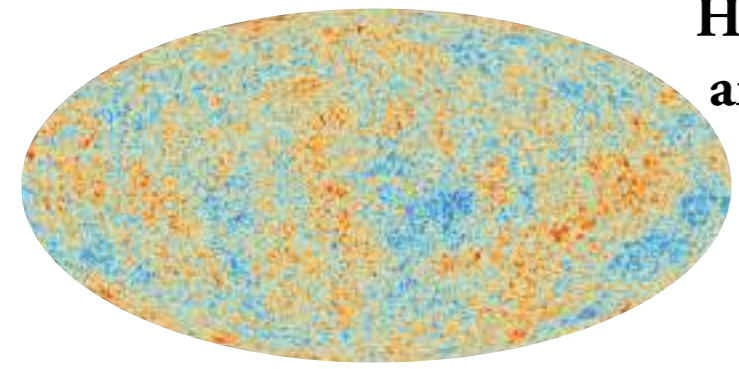
Why only δ ? → $\nabla \cdot v = -\dot{a}f\delta$

NB: only divergent (no tidal) but periodic boundaries

Standard cosmological simulations

Initial conditions (ICs)

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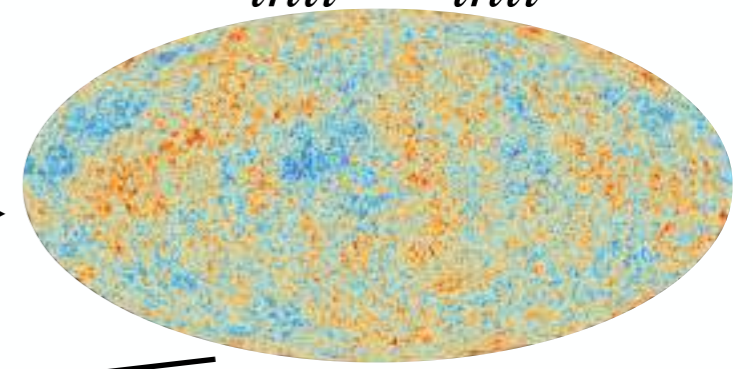
Homogeneous
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Universe

Gaussian
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field

$$P(k)$$

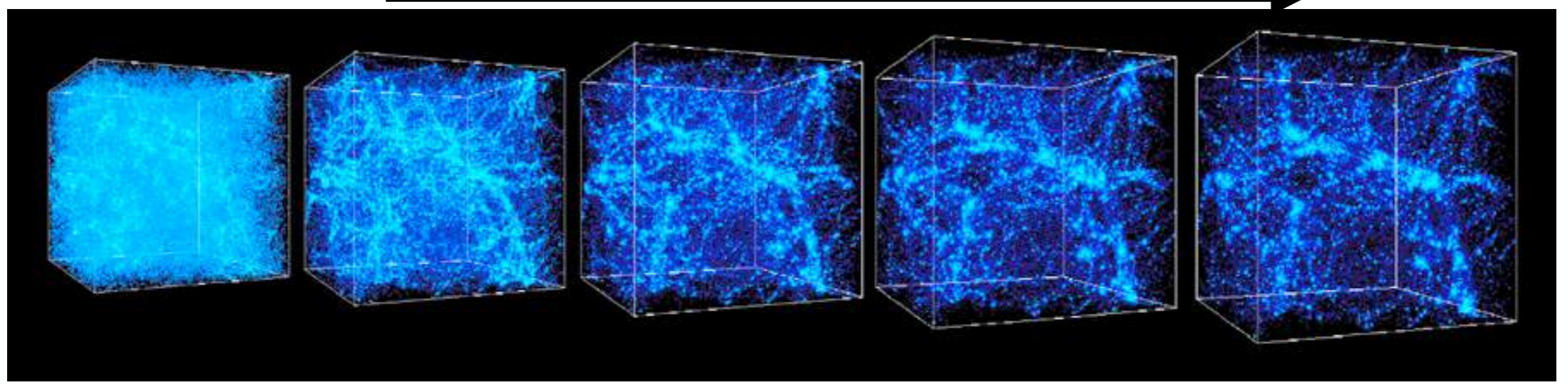
$$\delta(\mathbf{k}) = \sqrt{P(\mathbf{k})} \cdot \omega(\mathbf{k})$$

initial conditions of
a random patch of
the Universe
 $\{\delta_{init}, v_{init}\}$



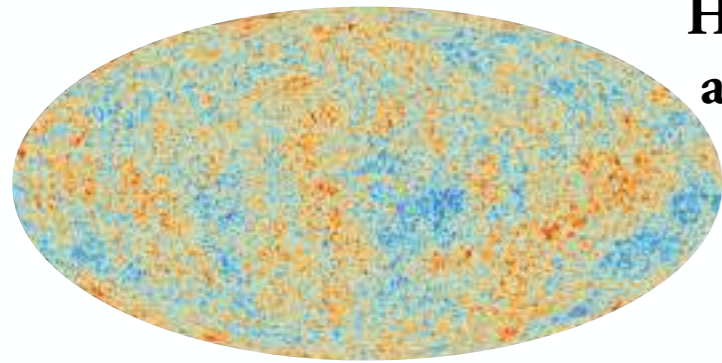
Evolution

Linear perturbation theory + "kick"



Constrained cosmological simulations

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Homogeneous
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$P(k)$

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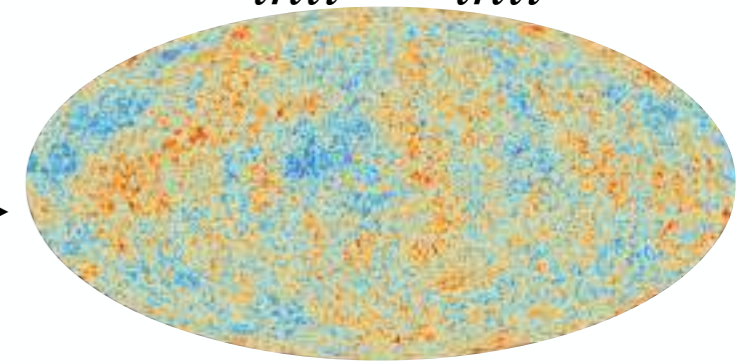
→

$$\delta(\mathbf{k}) = \sqrt{P(\mathbf{k})} \cdot \omega(\mathbf{k})$$

?

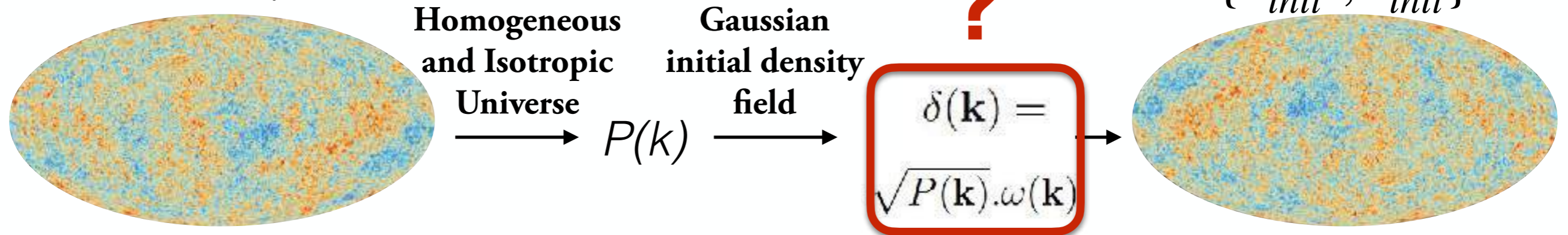
initial conditions of
the local Universe

$\{\delta_{init}, v_{init}\}$



Constrained cosmological simulations

Part of the Universe at
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Type of constraints

Redshift



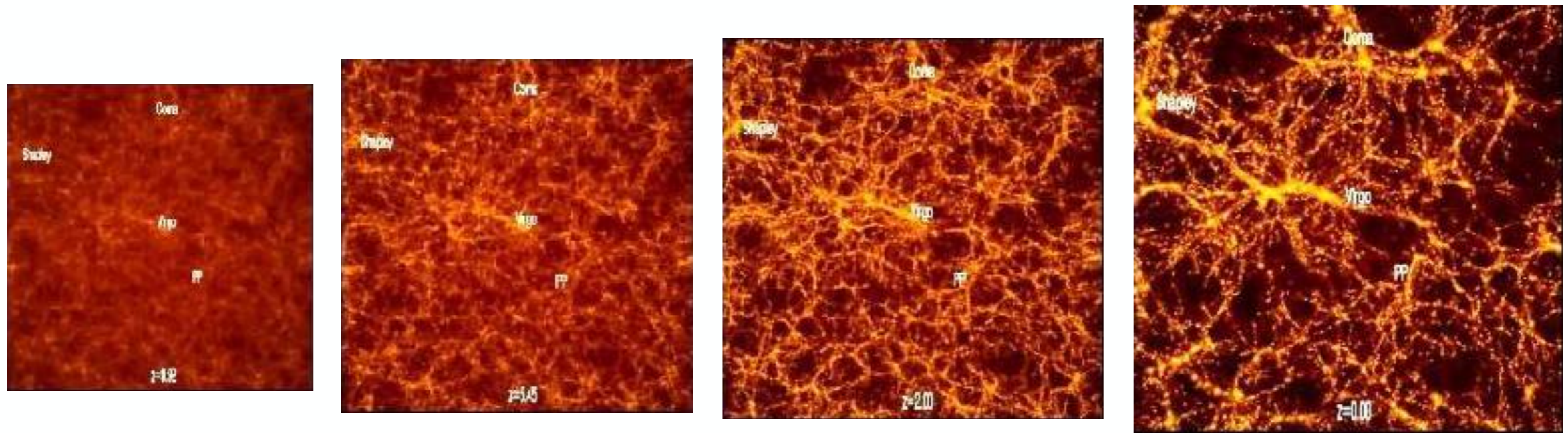
Peculiar velocity



NB: both with pros and cons!



Constrained cosmological simulations ▶ e.g. CLONES



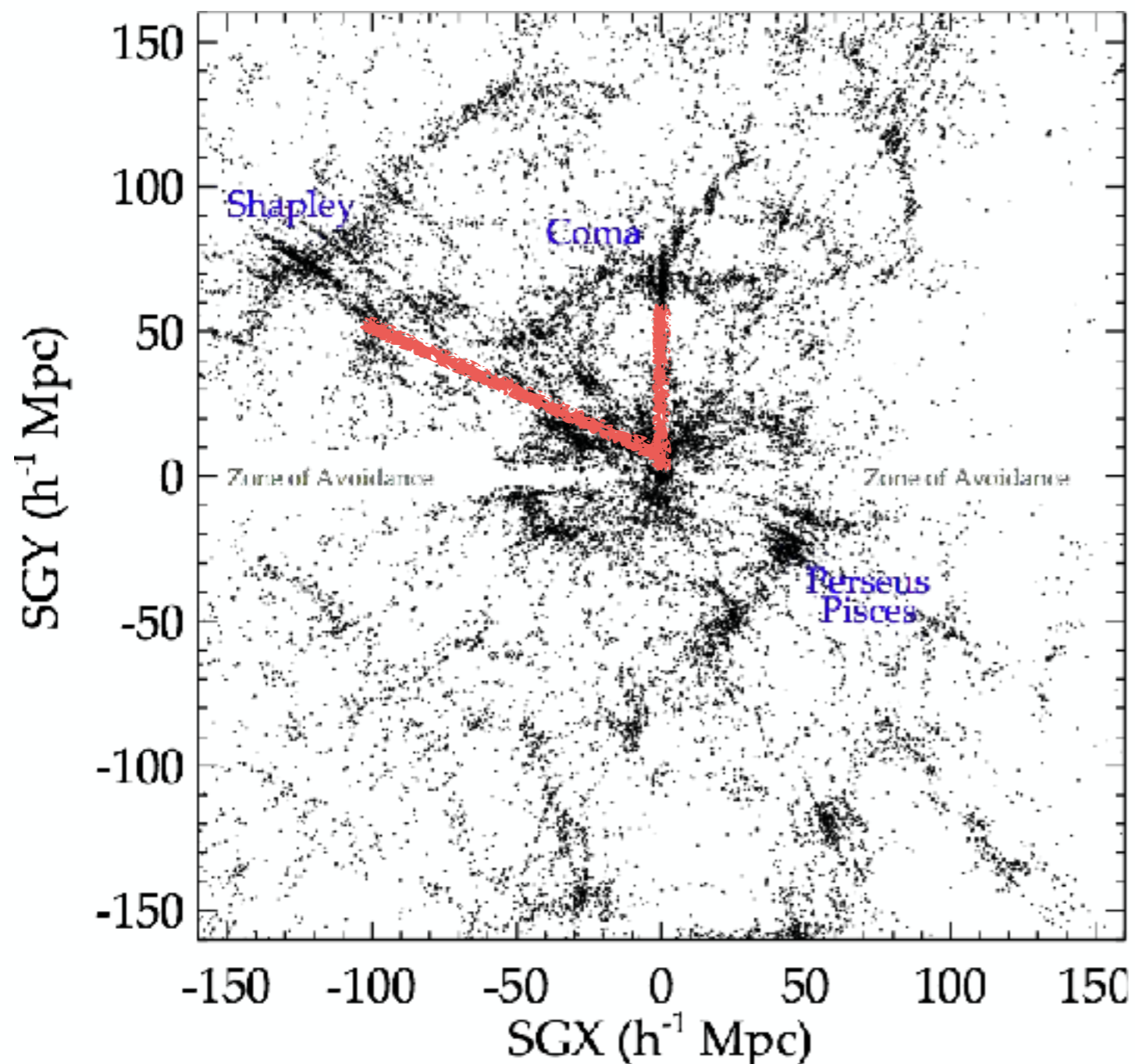
Evolution

Sorce+2016

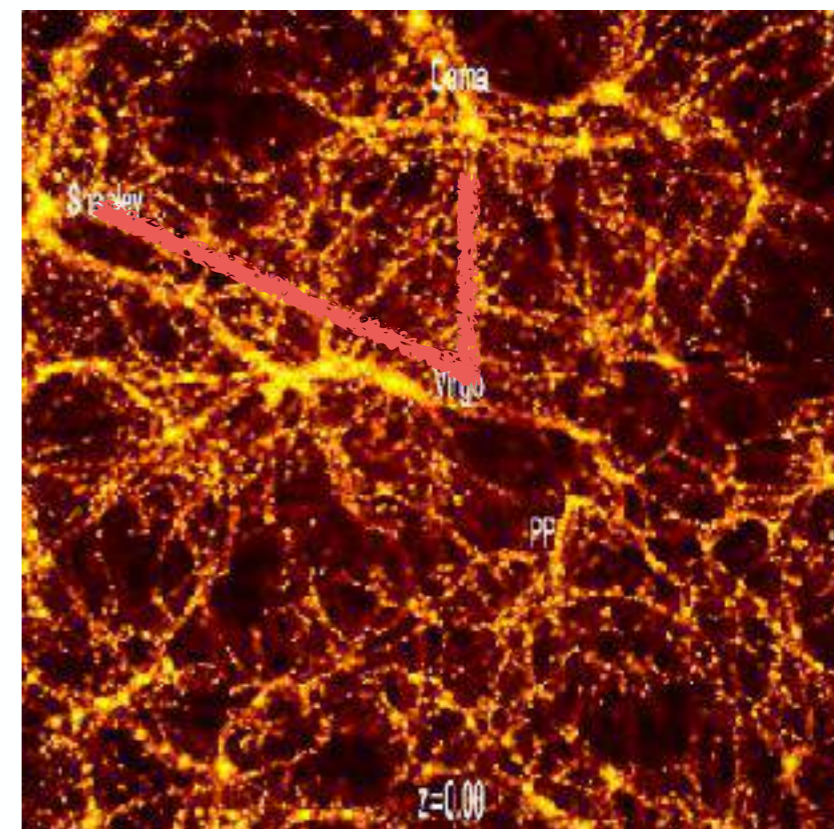
Sorce2018



CLONES = Constrained LOcal & Nesting Environment Simulations

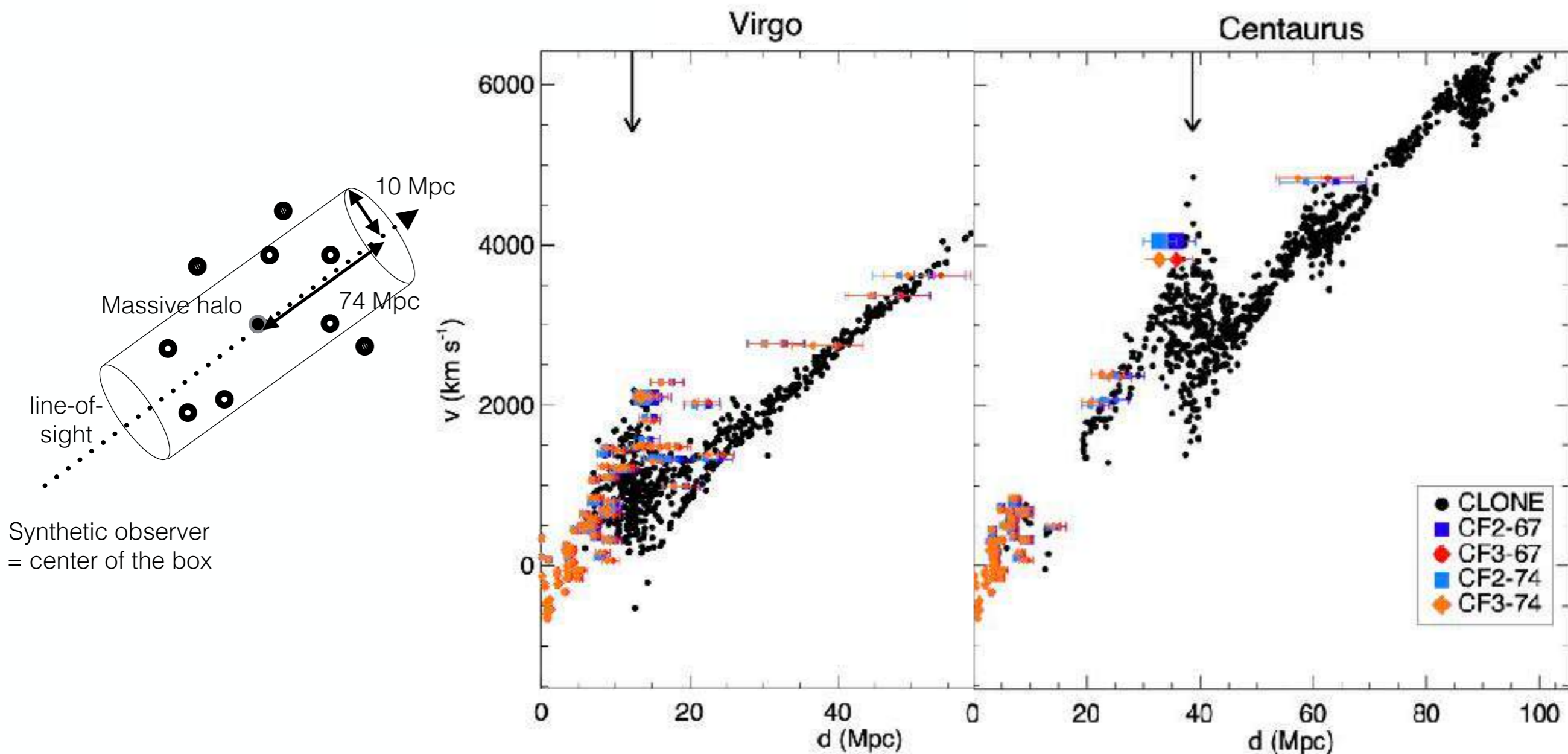


Note the fingers of gods



500 Mpc/h, 1024^3 particles,
DM only, Planck cosmology

Velocity wave signatures in the Hubble diagram



500 Mpc/h, 2048^3 particles, DM only, Planck cosmology



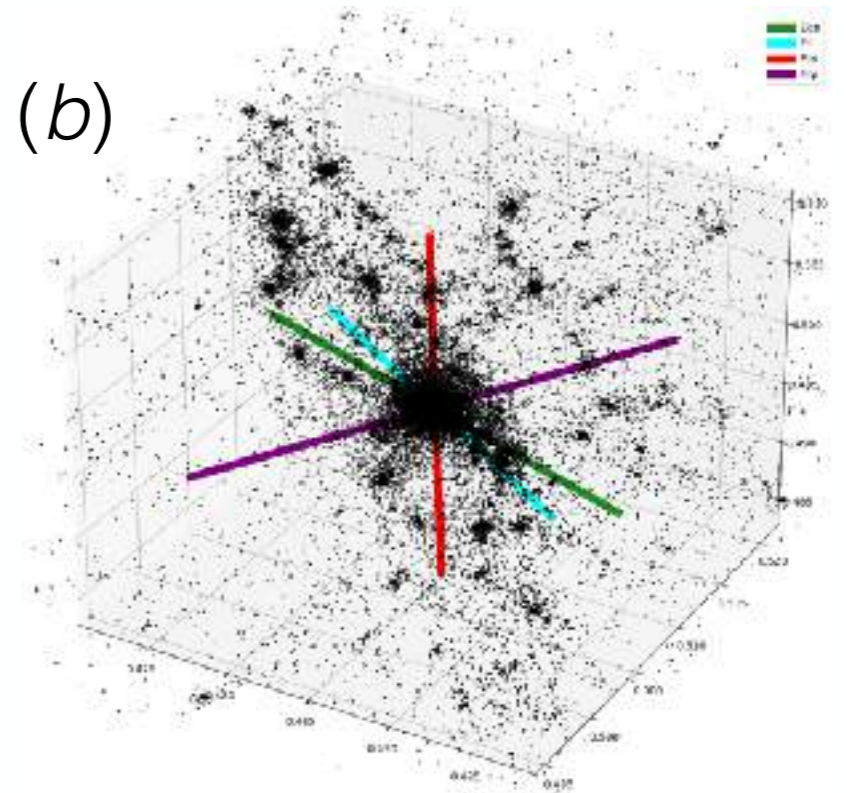
Théo Lebeau

Project: using a CLONE of the local Volume that contains replicas of local clusters to study the impact of

- the dynamical state of the cluster (substructures, morphology)
- the local Environment (connectivity)
- the formation history (accretion from filaments, merging)

on the hydrostatic mass bias (b)

-> *Example of the projection effects on the hydrostatic mass bias in the case of the Virgo cluster*

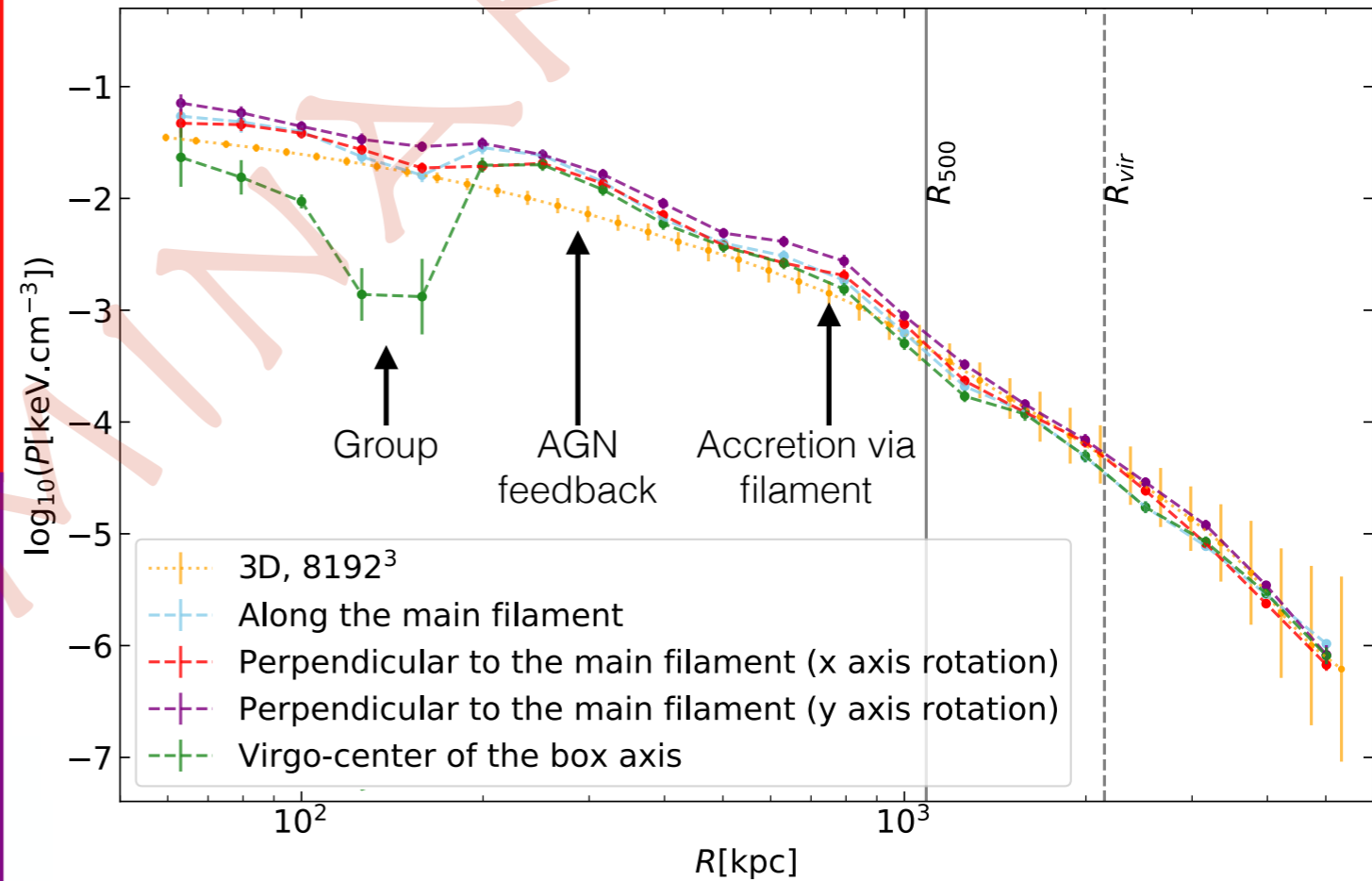
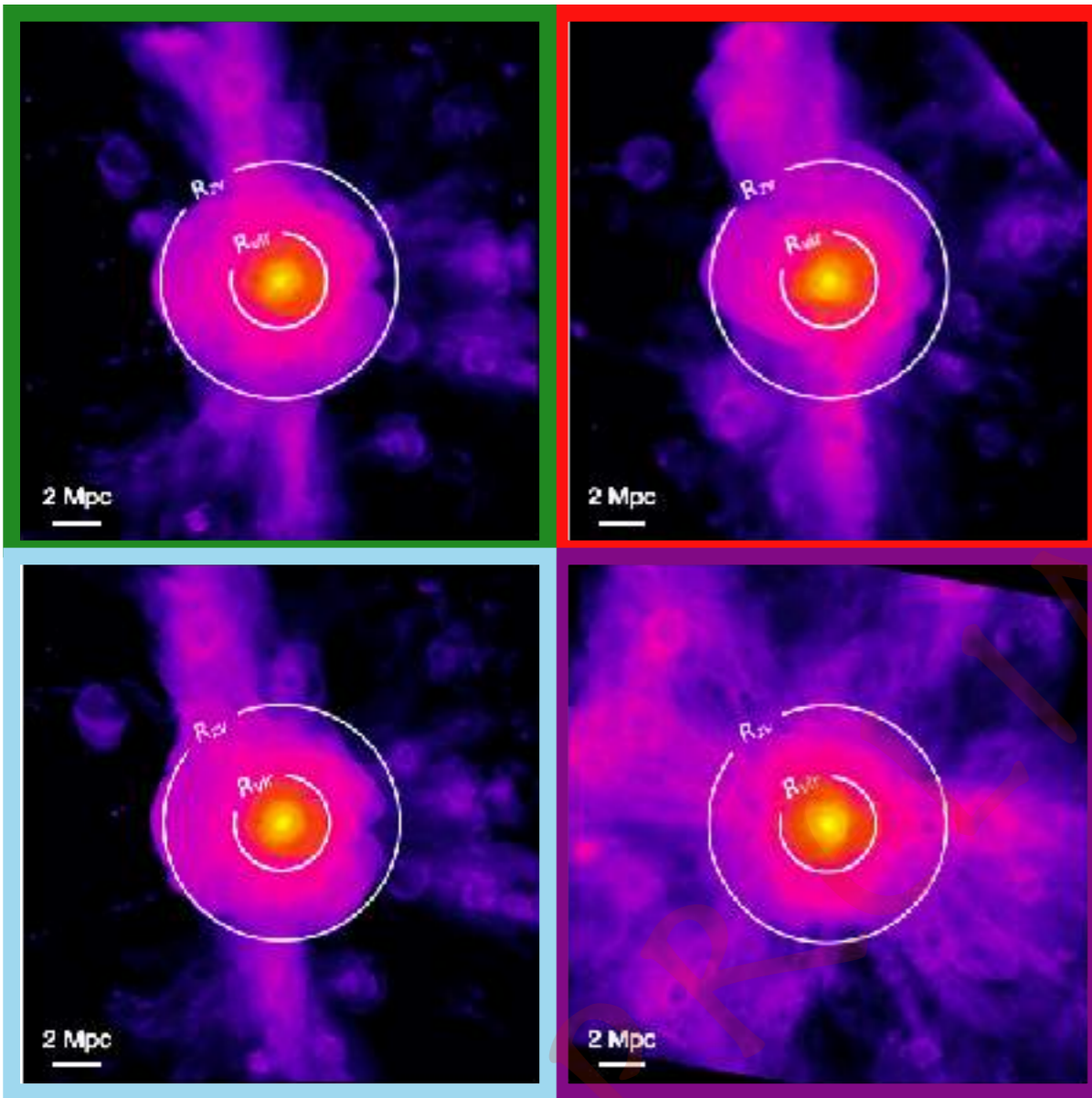


CLONES: hydrostatic mass bias & projection effects



Théo Lebeau

Example of the Virgo galaxy cluster



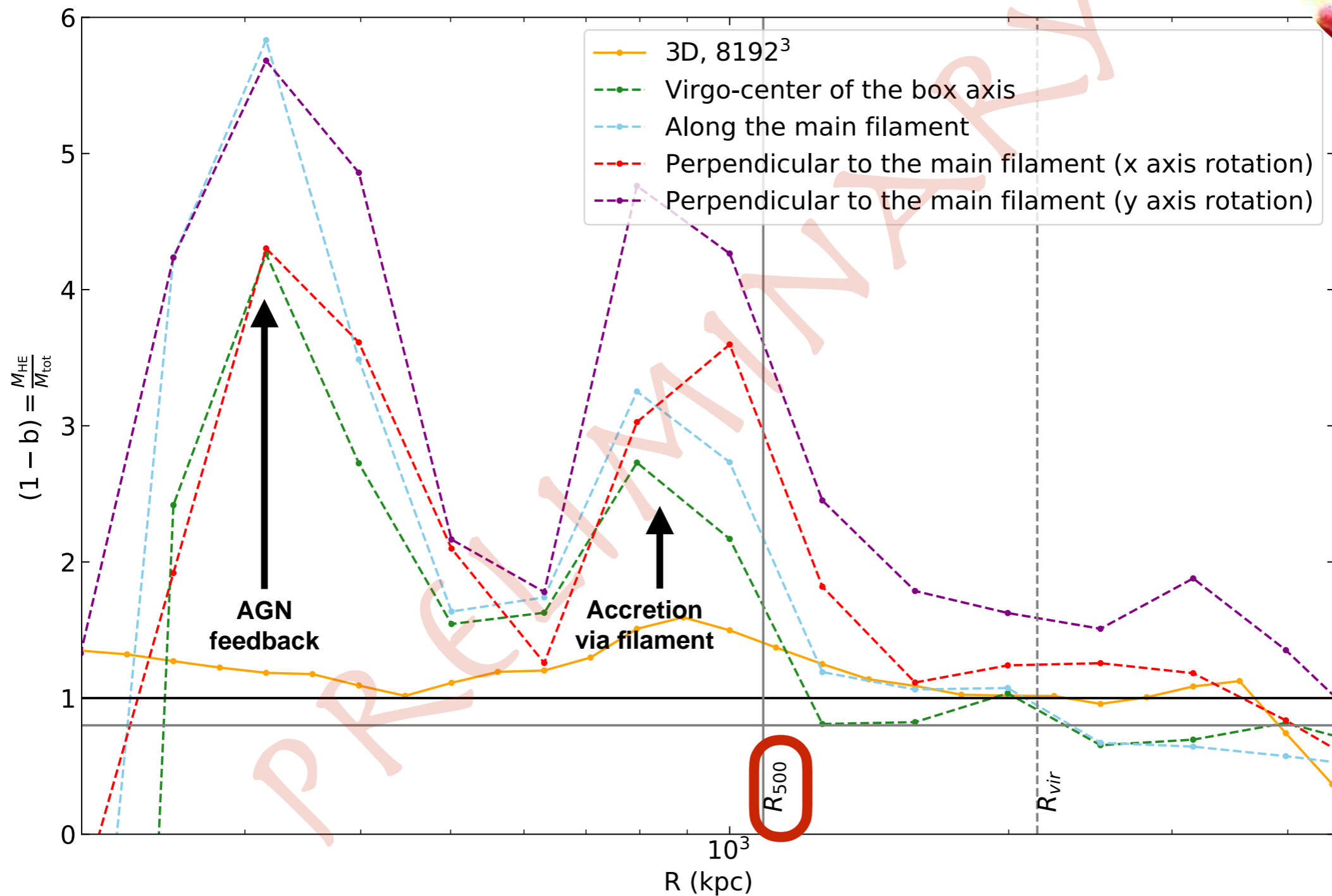
$$M_{HE}(r) = - \frac{r P_{th}(r)}{G \mu m_p n_e(r)} \frac{d \ln P_{th}(r)}{d \ln r} = (1 - b) M_{tot}$$

Lebeau+in prep.

CLONES: hydrostatic mass bias & projection effects



Théo Lebeau

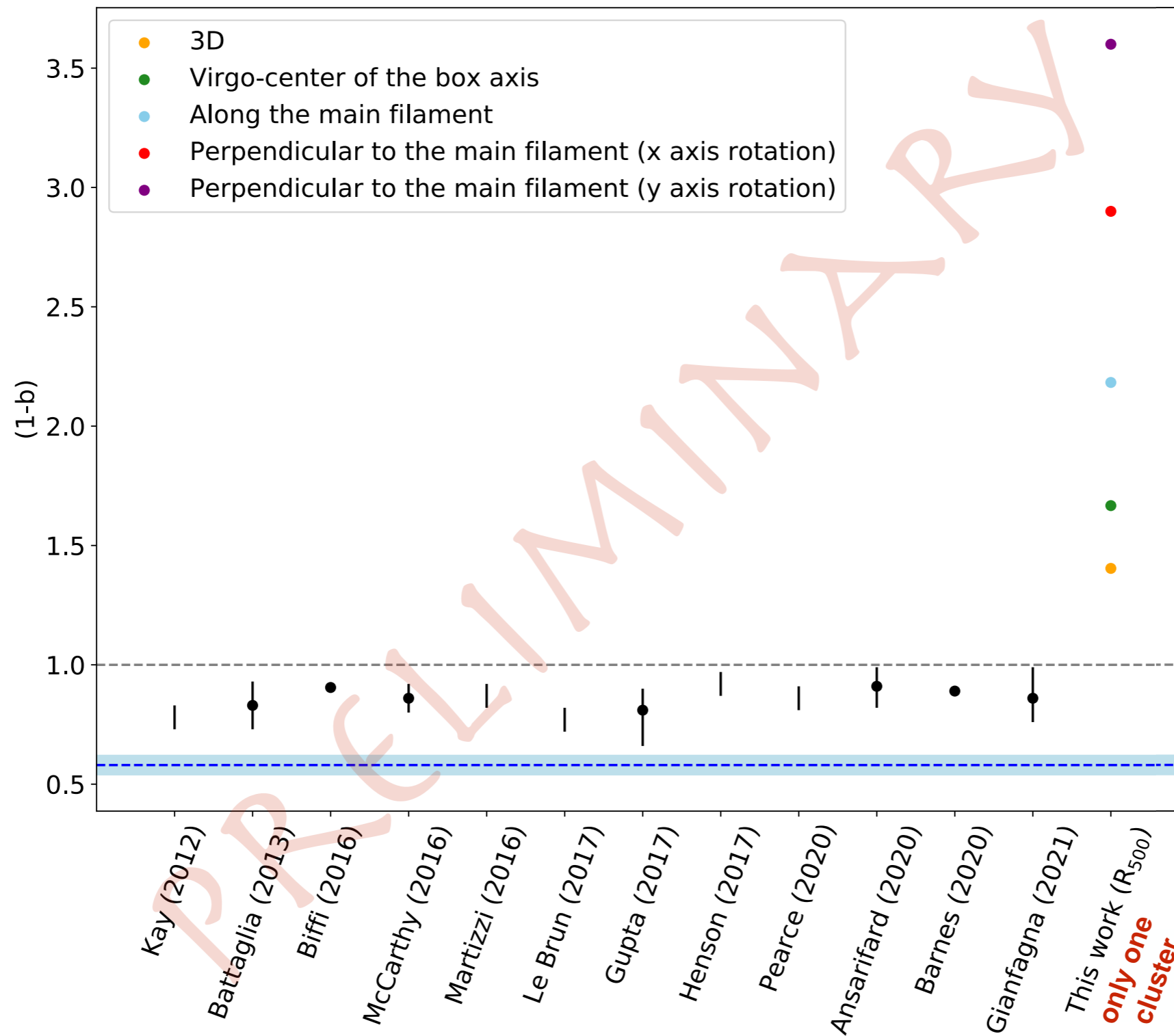


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CLONES: hydrostatic mass bias & projection effects



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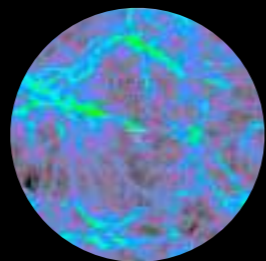


Lebeau+in prep.



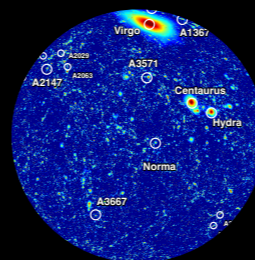
Virgo Cluster

(Sorce+2016, 2019, 2021, in prep., Olchanski & Sorce 2018, Lebeau+ in prep.)



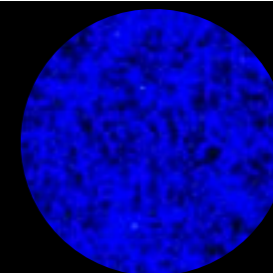
Cosmic Rays in the local Universe

(Hackstein+2018, Boess+in prep.)



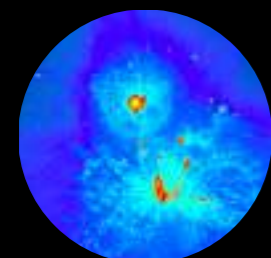
SLOW: local web

(Dolag, Sorce+2023)



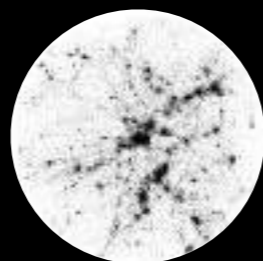
LOCALIZATION: local cluster signatures

(Sorce, Aghanim, Lebeau, Jung, Dolag)



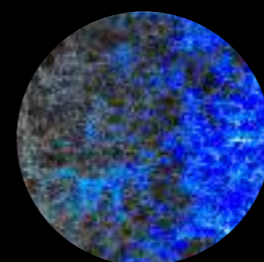
HESTIA: Local Group

Carlesi, Sorce+2016, Carlesi+2016, 2017, Libeskind+2020, Damle+2022, Newton+2022; Luis+2022, Dupuy+2022, Arora+2022, Khoperskov+2022a,b,c



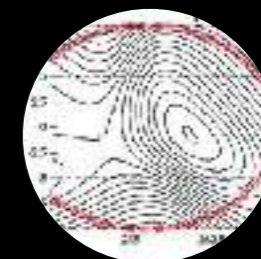
Coma connectivity

(Malavasi, Sorce, Dolag, Aghanim submitted)



CoDa: Reionization of the local Universe

(Ocvirk+2020, Lewis+2020, Gronke+2021, Sorce+2022, Lewis+2022, Park+2022)

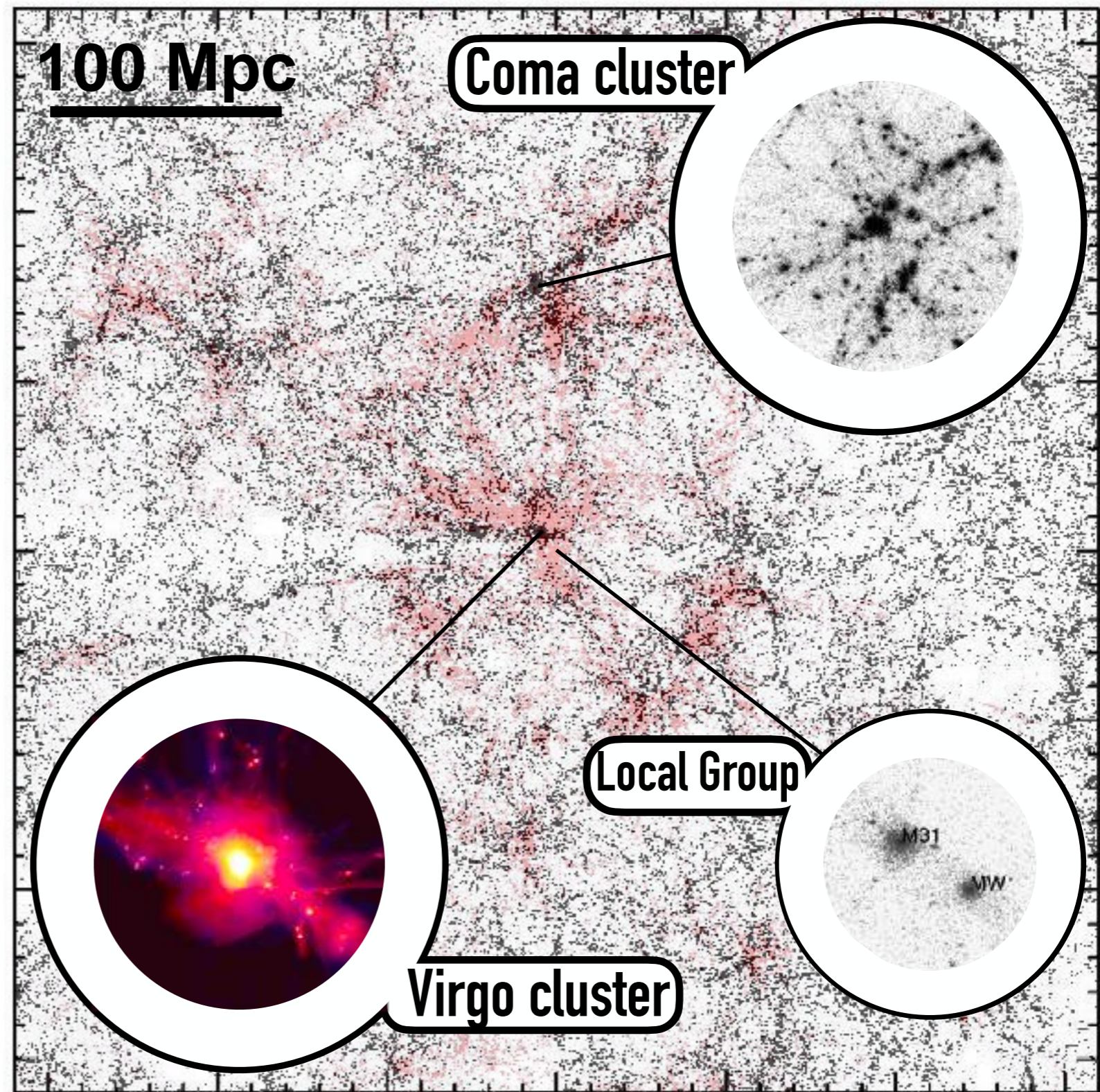


Zone of Avoidance

(Sorce+2017)

Conclusion

- **Standard** cosmological simulations give only the full uncertainty
- **Constrained** cosmological simulations can permit **reducing biases/systematics**
- **CLONES are constrained** cosmological simulations valid down to the cluster scales with induced smaller scales
- CLONES are **widely used** and **maybe you are the next users!**



**Thank you, Merci, Grazie,
Gracias, Danke,
Mahalo, 谢谢, ありがとう,
הודת, Obrigada, Dank u,
Tak, Cảm ơn, Dziękuję, 감사합니다
Kiitos, Aitäh, diolch, dankewol,
ಧನ್ಯವಾದಗಳು, ...***

* Missing your 'thanks' spelling? It means I did not get the chance to learn how to say it so far

